

# **HMS Cherenkov Calibration**

## **Method-1**

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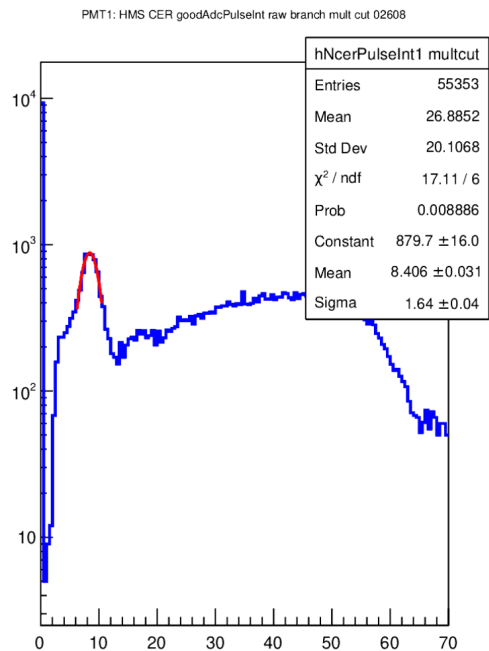
August, 2020

# Procedure

- As a clear SPE peak spectrum is seen in the HMS CER, the good ADC pulse integral is fitted with Gaussian function. Then the mean is extracted from the fit and we put  $1/\text{mean}$  in the calibration param.
- The cuts that were used on the ADC pulse integral (H.cer.goodAdcPulseInt) are the following:
  1. Requiring goodAdcmultiplicity == 1 for the PMT under calibration.
  2. Detector timing window cuts (H.cer.goodAdcTdcDiffTime) (already set in the corresponding parameter file in hallc\_replay ).

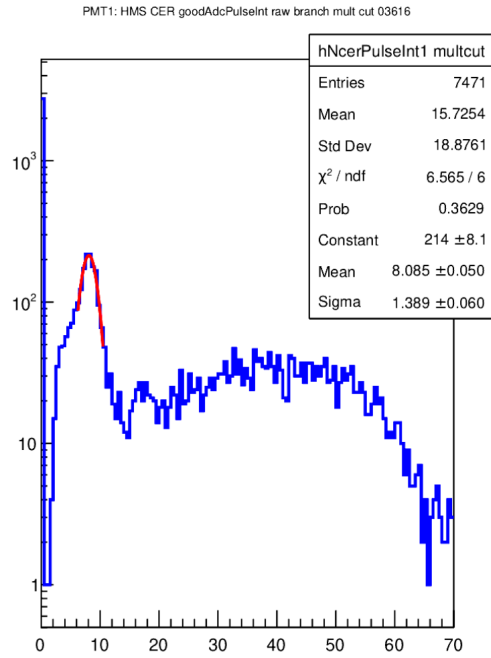
# PMT 1

- Gaussian fit to H.cer.goodAdcPulseInt (multiplicity==1) for different runs.



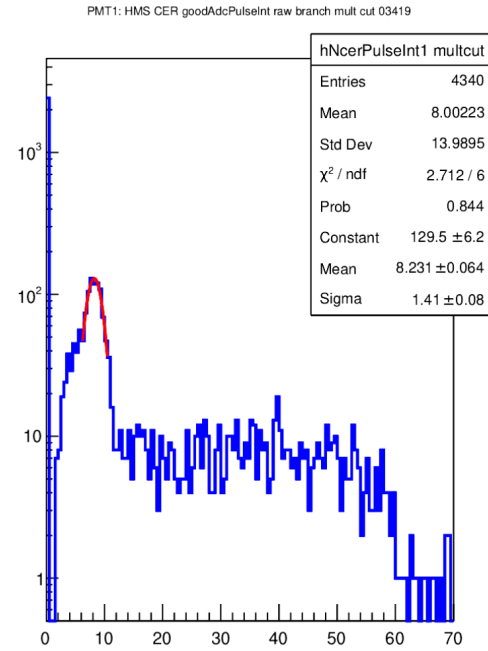
HMS Run 02608: 1-pass elastic, EL-CLEAN

$$1 / \text{mean} = 1 / 8.40605$$



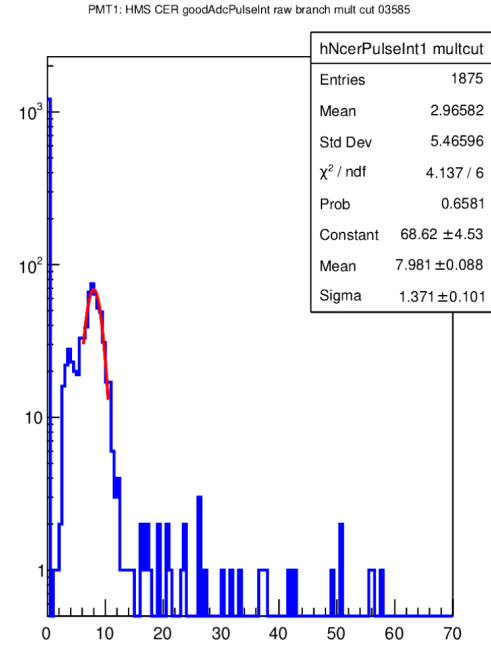
HMS Run 03616: C-Optics/Sieve, EL-REAL

$$1 / \text{mean} = 1 / 8.08504$$



HMS Run 03419: DIS, 3/4

$$1 / \text{mean} = 1 / 8.23131$$



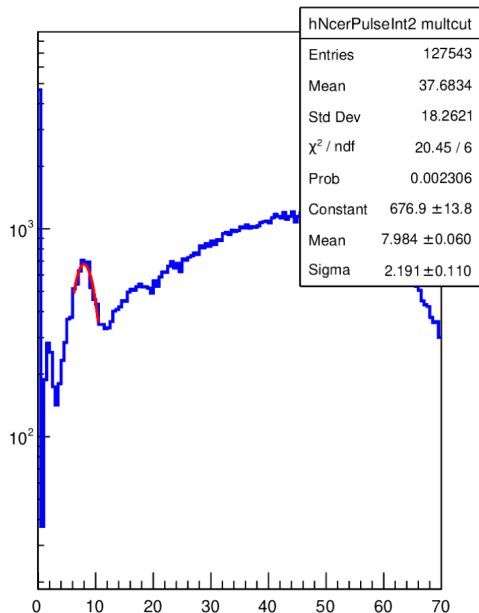
HMS Run 03585: Cosmics

$$1 / \text{mean} = 1 / 7.98096$$

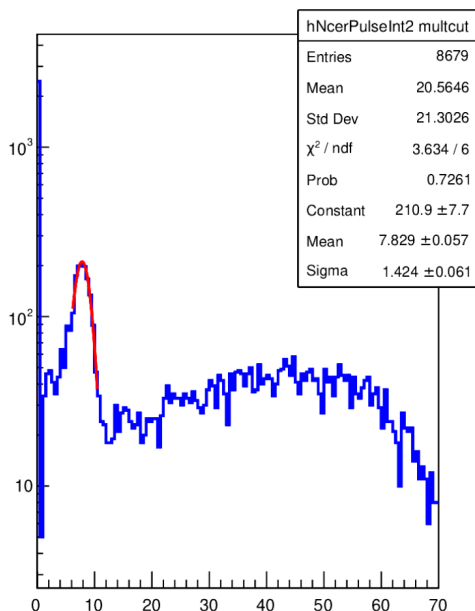
# PMT 2

- Gaussian fit to H.cer.goodAdcPulseInt (multiplicity==1) for different runs.

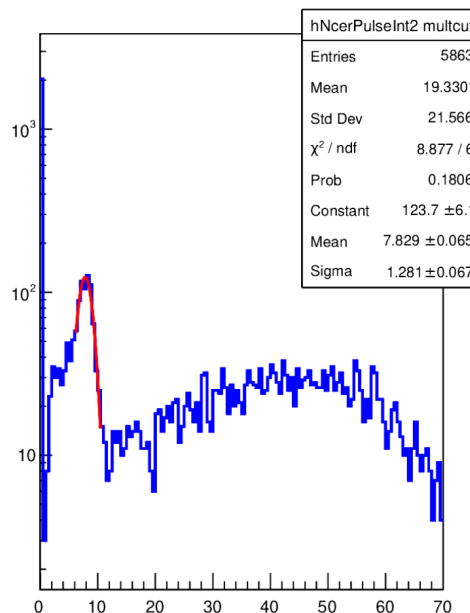
PMT2: HMS CER goodAdcPulseInt raw branch mult cut 02608



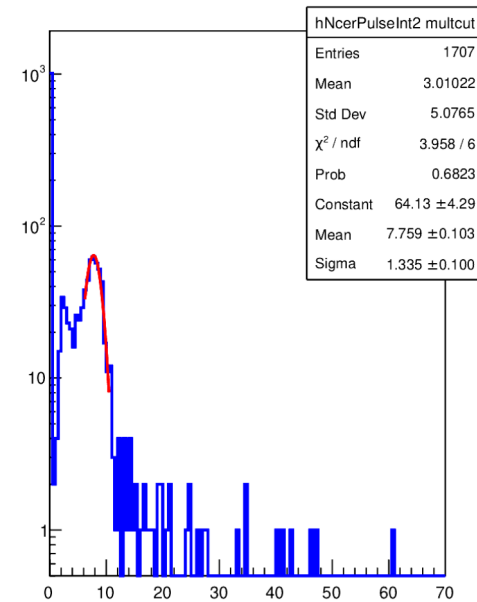
PMT2: HMS CER goodAdcPulseInt raw branch mult cut 03616



PMT2: HMS CER goodAdcPulseInt raw branch mult cut 03419



PMT2: HMS CER goodAdcPulseInt raw branch mult cut 03585



HMS Run 02608: 1-pass elastic, EL-CLEAN

$$1 / \text{mean} = 1 / 7.98363$$

HMS Run 03616: C-Optics/Sieve, EL-REAL

$$1 / \text{mean} = 1 / 7.82941$$

HMS Run 03419: DIS, 3/4

$$1 / \text{mean} = 1 / 7.8286$$

HMS Run 03585: Cosmics

$$1 / \text{mean} = 1 / 7.75936$$

# Calibration constants

The estimated calibration constants values per PMT are the following:

HMS Runs	PMT 1	PMT 2
Old Param file (Cosmics 2504)	1 / 8.409	1 / 7.735
2608 (Elastic) (Old replay)	1 / 8.40605	1 / 7.98363
3616 (C-Optics / Sieve) (Old replay)	1 / 8.08504	1 / 7.82941
3419 (DIS) (Old replay)	1 / 8.23131	1 / 7.8286
3585 (Cosmics) (Old replay)	1 / 7.98096	1 / 7.75936

# Questions

1. The calibration constants in the old param file were estimated from Cosmic run 02504.

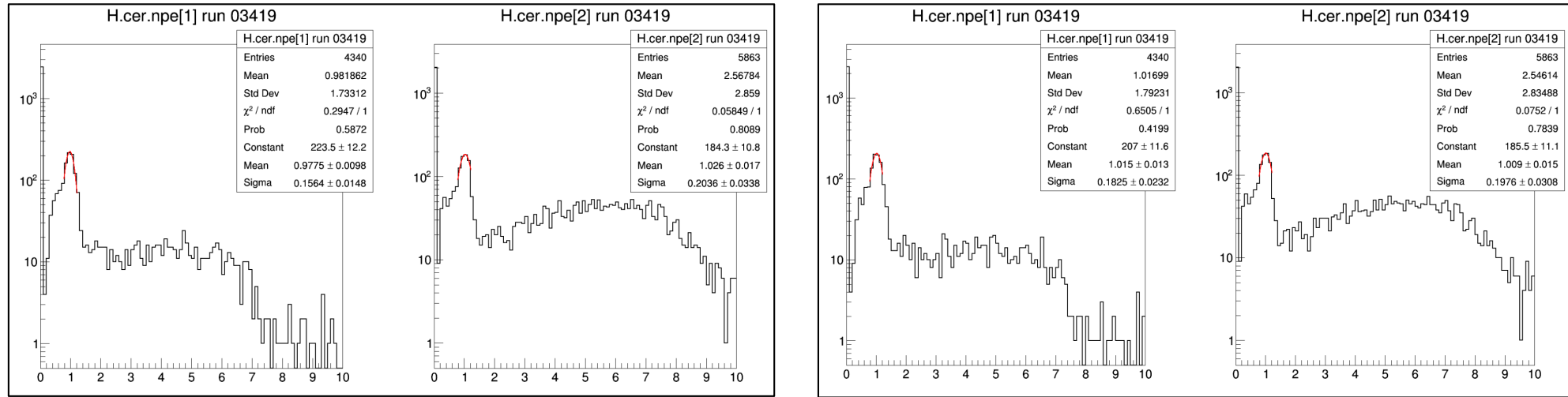
I have done the calibration with a new cosmic run 03585. There is a slight difference in the fit parameters for the two cosmics.

Can we expect that ?

2. Do we need to apply anymore cuts (eg. Mirror cuts etc.) like method 2 ?

3. The calibration constants are a bit different for different runs. How should we choose one ?

# New replay: New param file with PMT 1= 1 / 8.08504 and PMT 2 = 1 / 7.82941 (obtained from run 2608 SPE fit)



HMS run 3419 (DIS): Comparing H.cer.npe for old and new replay

H.cer.npe is fitted with Gaussian function. The mean of the fits are the following:

- **Old Replay:** PMT1 =  $0.9775 \pm 0.0098$ , PMT2 =  $1.026 \pm 0.017$
- **New Replay:** PMT1 =  $1.015 \pm 0.013$ , PMT2 =  $1.009 \pm 0.015$

**NPE has changed a little!**